



# Microchannel plate single photon imaging detectors for astronomy

*Do they need to be replaced?*

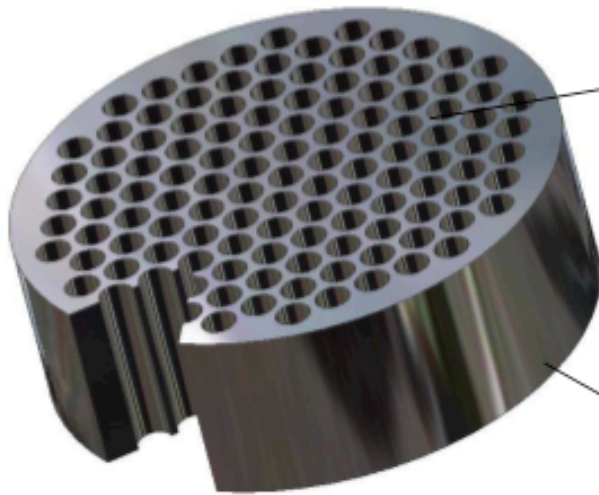
*John Vallerga  
Jason McPhate, Anton Tremis & Oswald Siegmund  
Space Sciences Laboratory  
University of California, Berkeley*



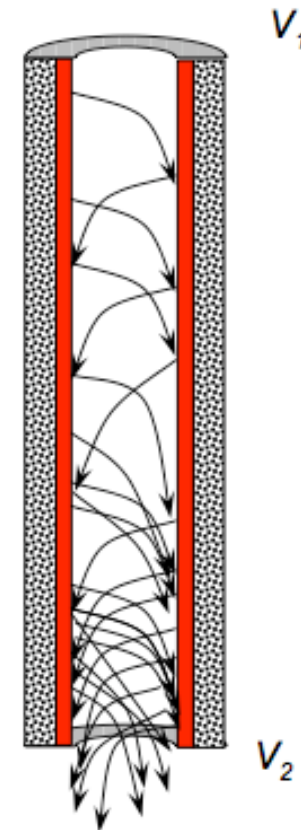
# Microchannel Plate

*electron amplifier that retains position information*

**Circular-pore MCP**



**Single pore**



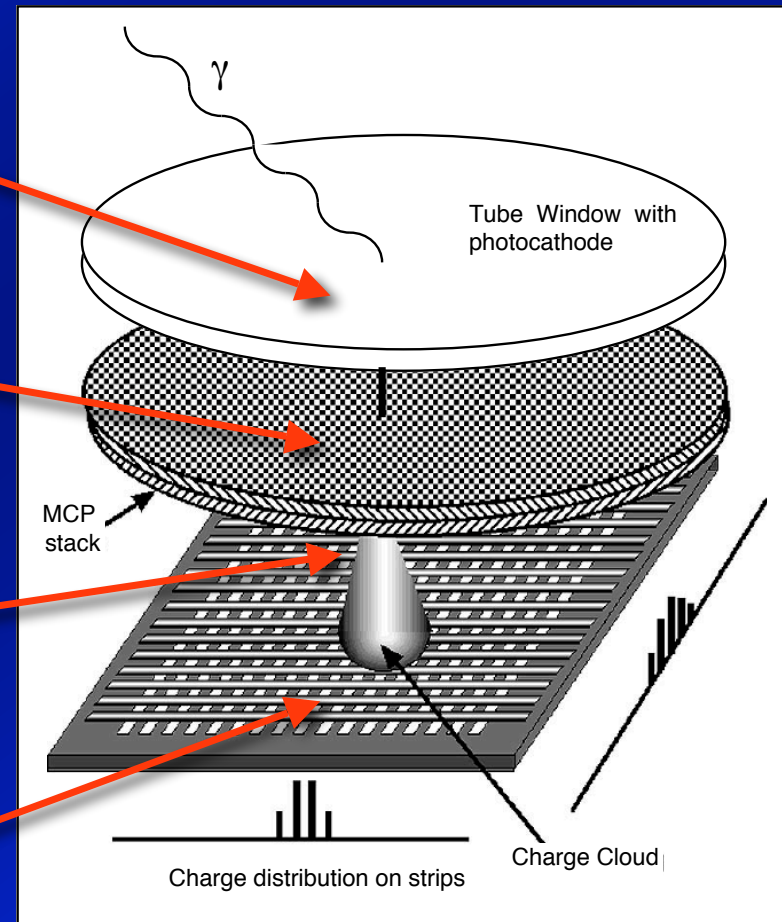
*Pore size*       $\sim 2 \text{ to } 40\mu\text{m}$   
*Gain*            $\sim 10 \text{ to } 10^8$   
*Time Jitter*     $\sim 30 \text{ ps}$



# Microchannel Plate Detector

*Photon counting, imaging, with event time tagging*

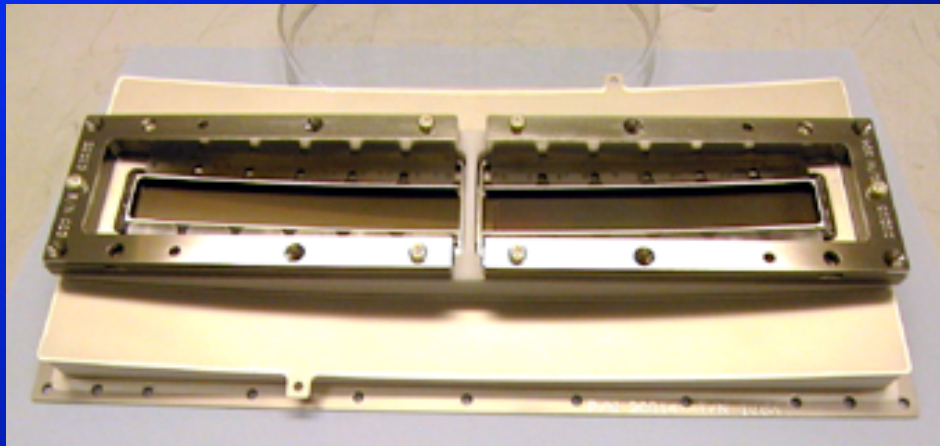
- Photocathode converts photon to electron
- MCP(s) amplify electron by  $10^4$  to  $10^7$
- Rear field accelerates electrons to anode
- Patterned anode measures charge centroid





# MCP Detectors at SSL Berkeley

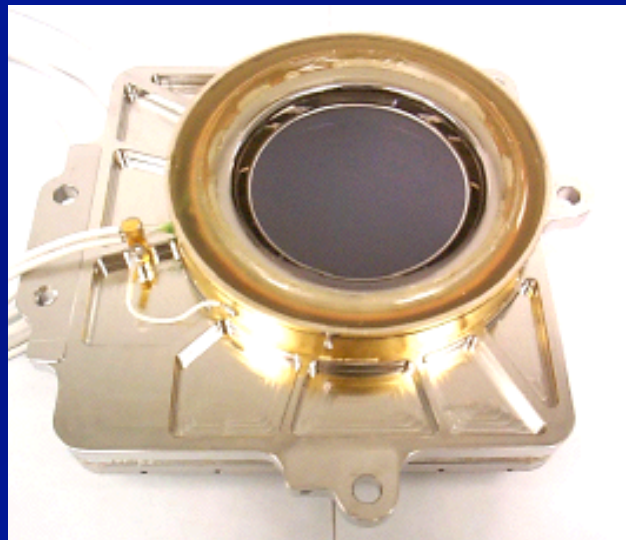
COS FUV for Hubble (200 x 10 mm windowless)



18 mm Optical Tube



GALEX  
68 mm NUV  
Tube

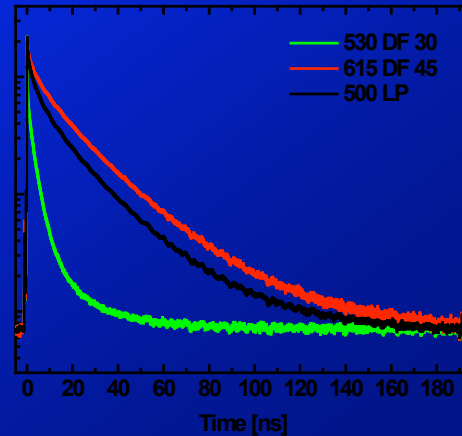
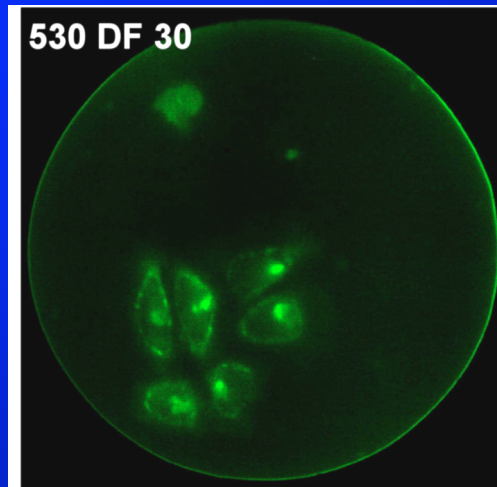


*200+ “detector years” in  
space including mission to  
Pluto*



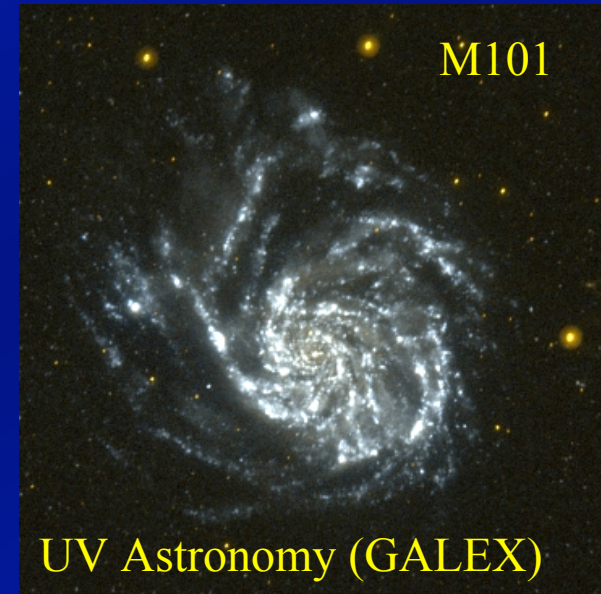


# Microchannel Plate Sensor Applications



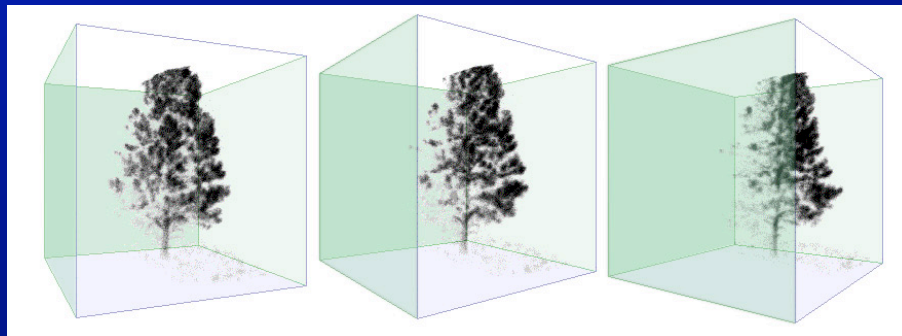
Fluorescent dye decay time

Biological lifetime fluorescence imaging (UCLA)



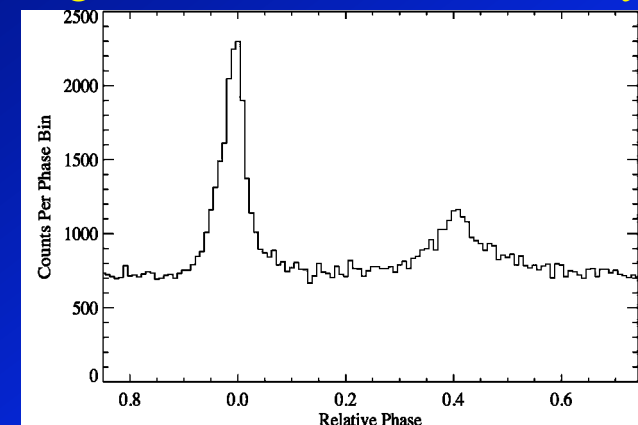
UV Astronomy (GALEX)

Low Light Ladar - 3D imaging



Ponderosa Pine 3D data cube - RULLI, LANL

High time resolution Astronomy

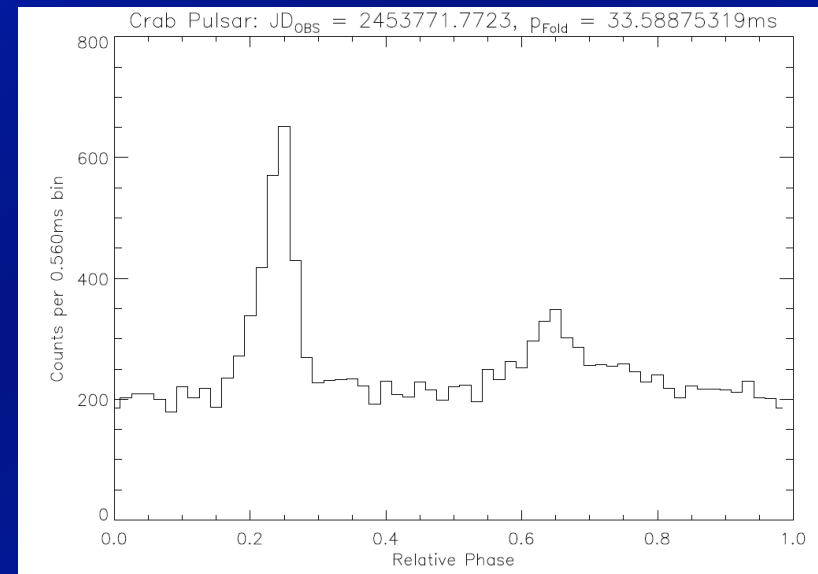
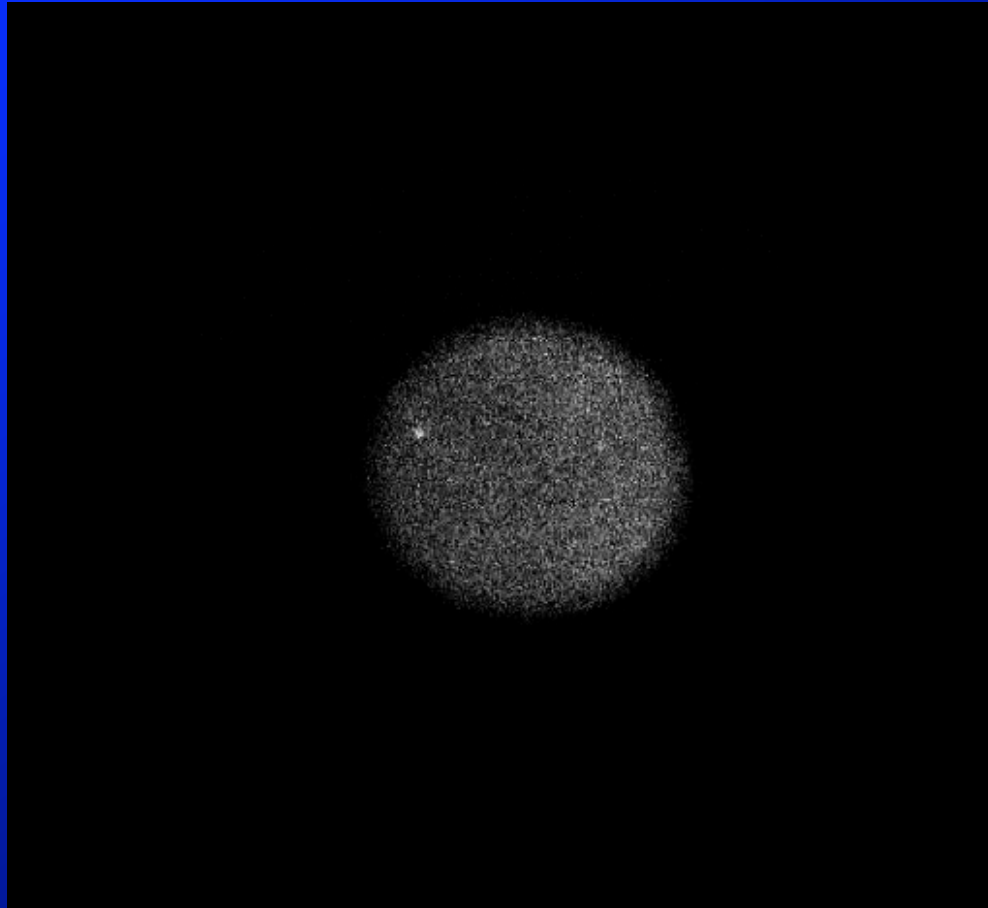


33ms period, 250 $\mu$ s resolution

KISS Workshop: Single Photon Counting Detectors, John Valleja, [john.valleja@berkeley.edu](mailto:john.valleja@berkeley.edu)



# High Speed Imaging Astronomy



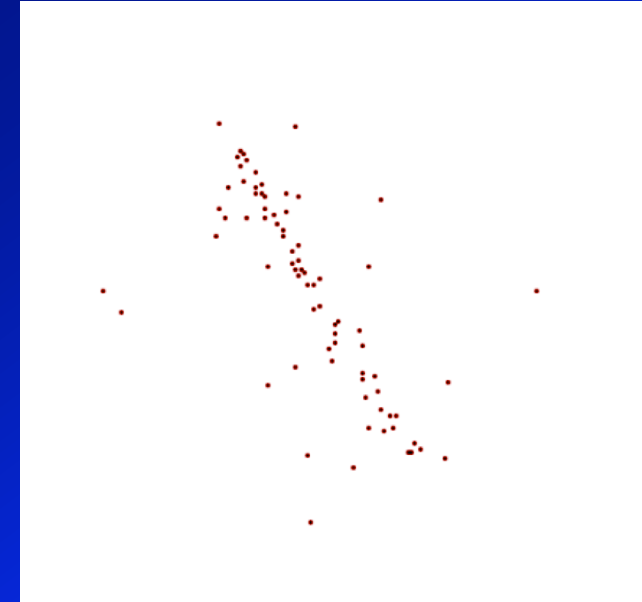
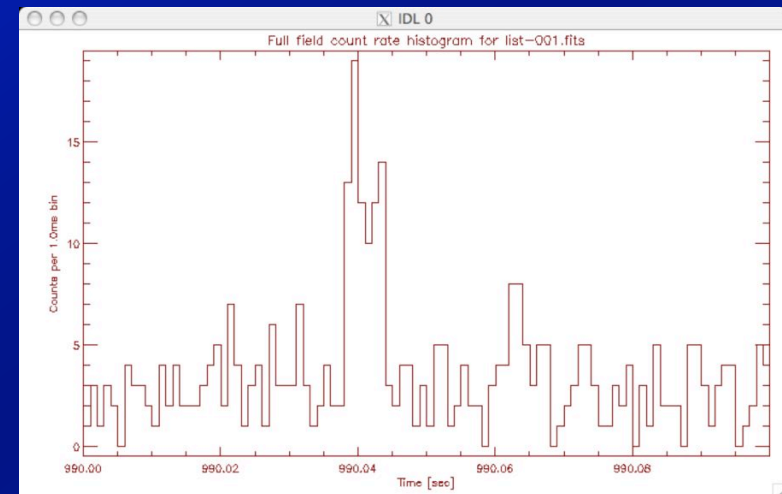
*Crab Pulsar B band light curve  
(1 m telescope on Moonlit night)*



# Meteor & Satellite Detection

Meteor track in AH-Her field

5 ms to cross 6' FOV.  
~11 km/sec if at 30 km

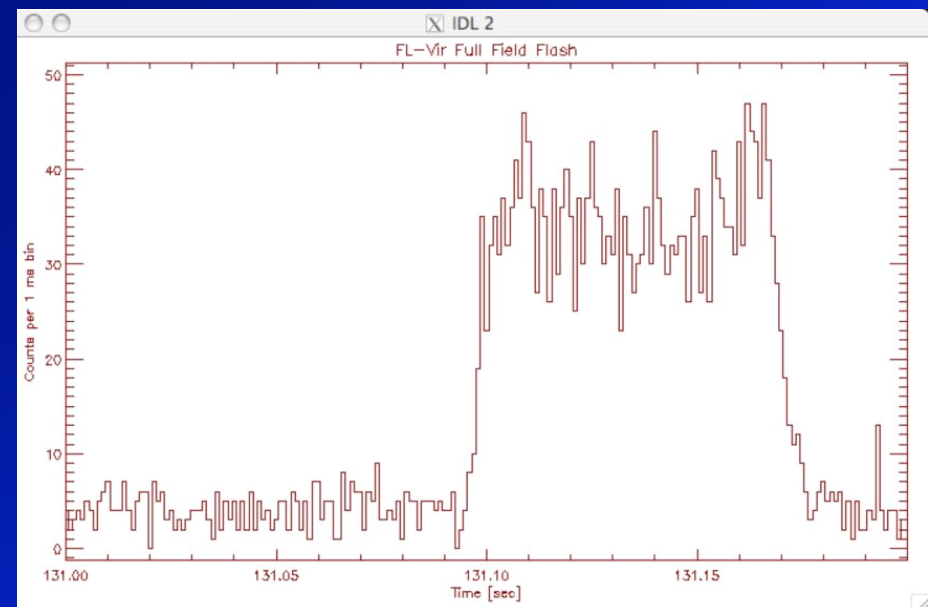




# Meteor & Satellite Detection



Satellite track in FL-Vir field  
~8.5 km/sec in low earth orbit







# Photon Counting MCP Detectors

## Advantages

<i>Spatial resolution</i>	$< 12\mu\text{m FWHM}$
<i>Temporal resolution</i>	$< 100\text{ ps}$
<i>Format</i>	$> 100\text{ mm}$
<i>Dynamic range</i>	$> 10^8$
<i>Radiation hard</i>	
<i>Curved focal planes</i>	
<i>Room Temperature</i>	

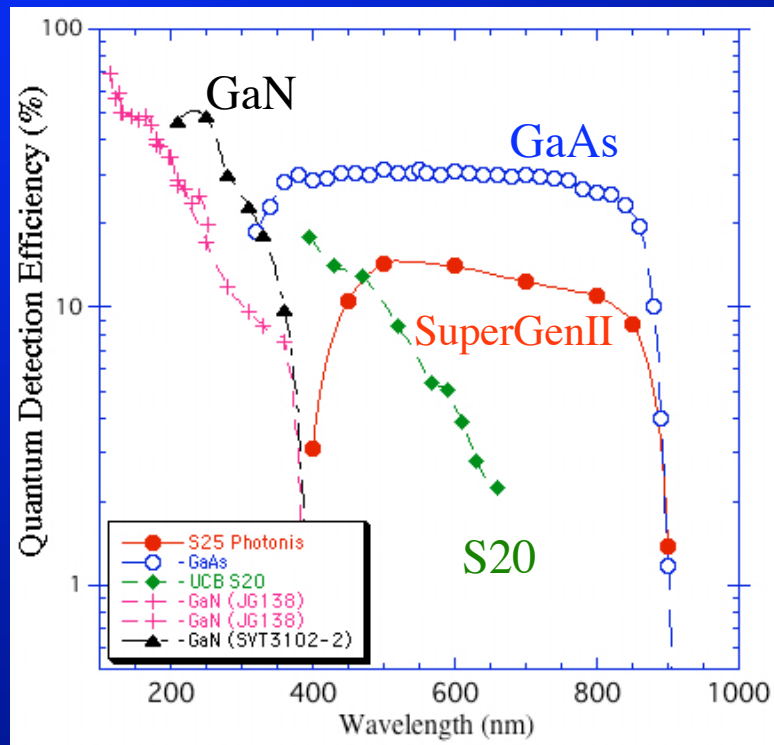
## Disadvantages

<i>Vacuum operation</i>	<i>pumps, tubes or space</i>
<i>Lifetime</i>	$\sim \text{Coulombs cm}^{-2}$
<i>Photoelectric QE</i>	$\sim 50\% \text{ to } 10^{-12}$



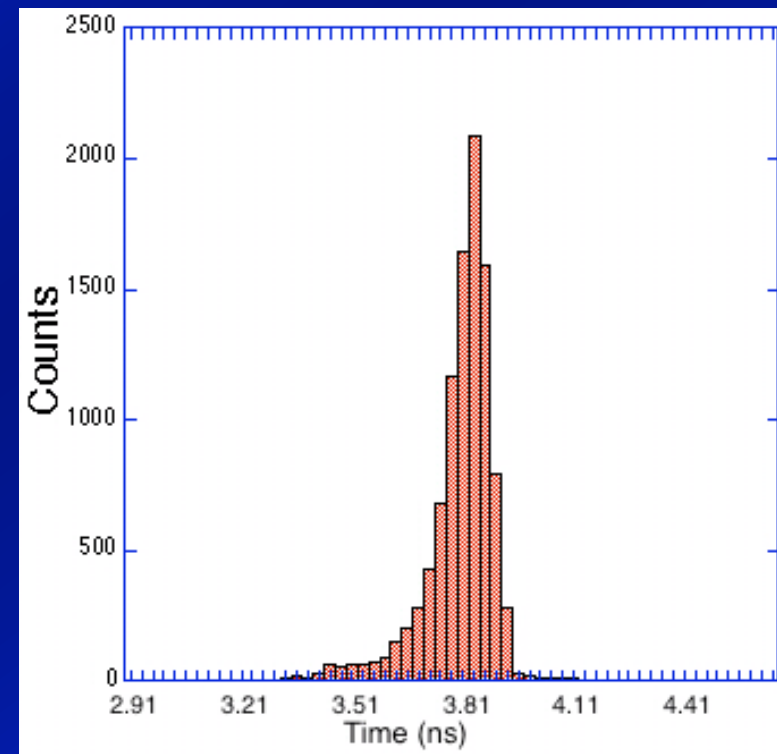
# Performance examples

## Photocathode QE



Photocathode types for UV, visible & NIR are improving.

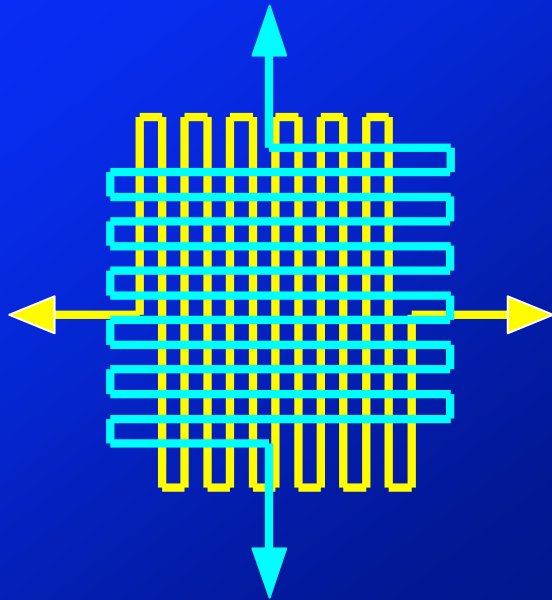
## Temporal resolution



MCP timing jitter ~100 ps FWHM measured with pulsed laser

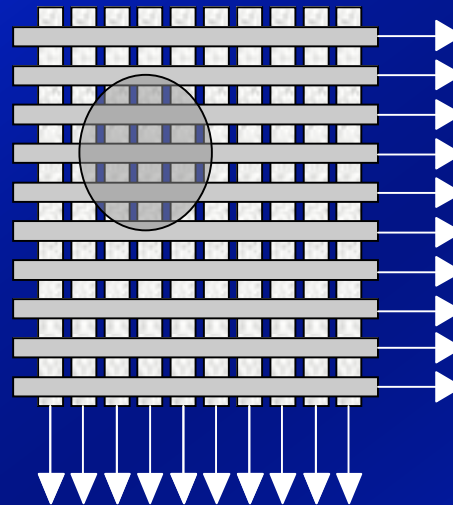


# Readout Anode Types (partial list)



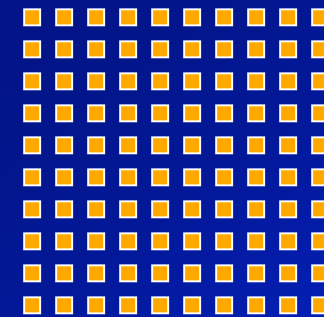
Cross Delayline  
(XDL)

4 amps  
Gain  $\sim 10^7$   
Rate  $\sim 200\text{kHz}$   
 $\Delta t \sim 100\text{ps}$



Cross Strip  
(XS)

2 x N amps  
Gain  $\sim 10^6$   
Rate  $\sim 2\text{MHz}$   
 $\Delta t \sim 100\text{ps}$



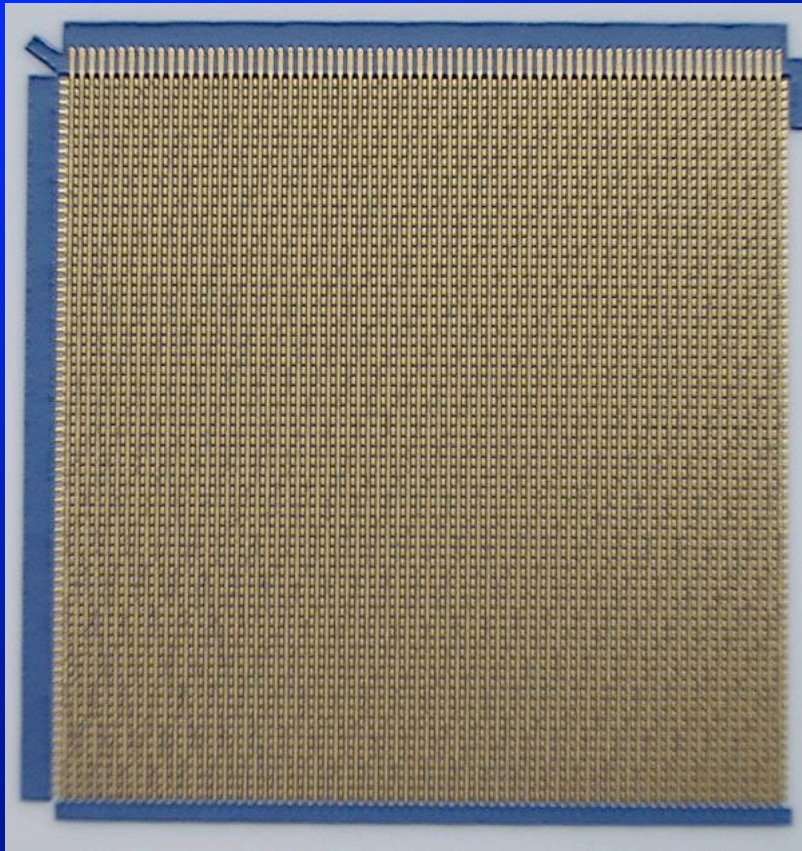
Medipix ASIC  
Intensified CCD

N x N amps  
Gain  $\sim 10^4$   
Rate  $\sim 200\text{MHz}$   
 $\Delta t \sim 1\text{ms}$

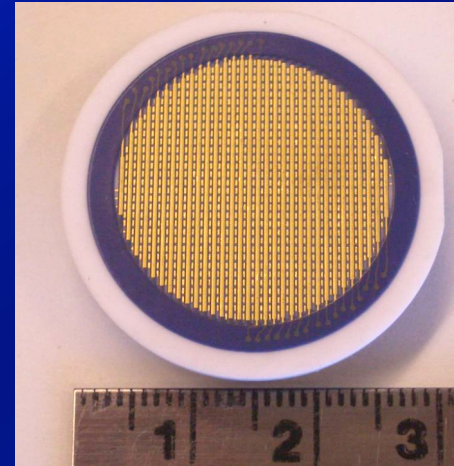




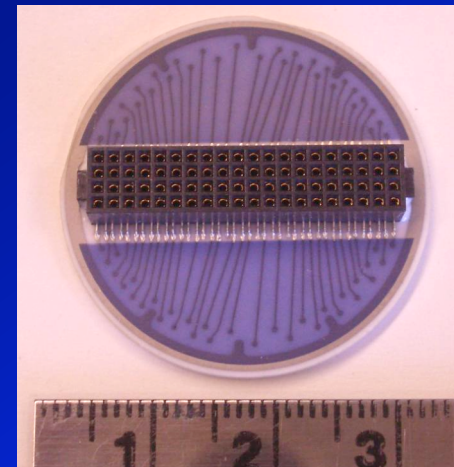
# *Cross Strip Anode Designs*



45mm square Cross Strip Anode  
with 0.64 mm finger period. All metal  
and ceramic.



22mm round cross strip

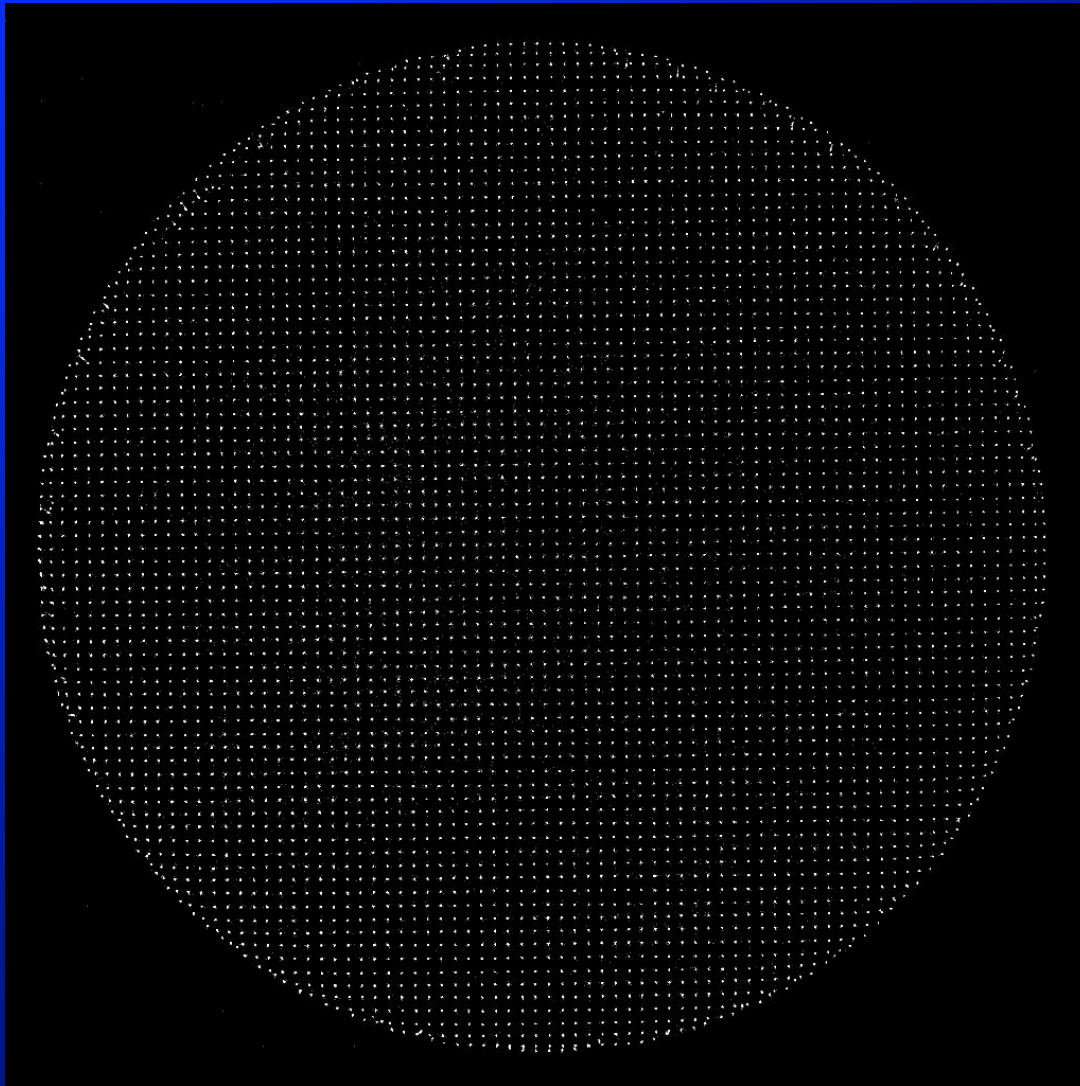


22mm round cross strip  
showing vias and connector

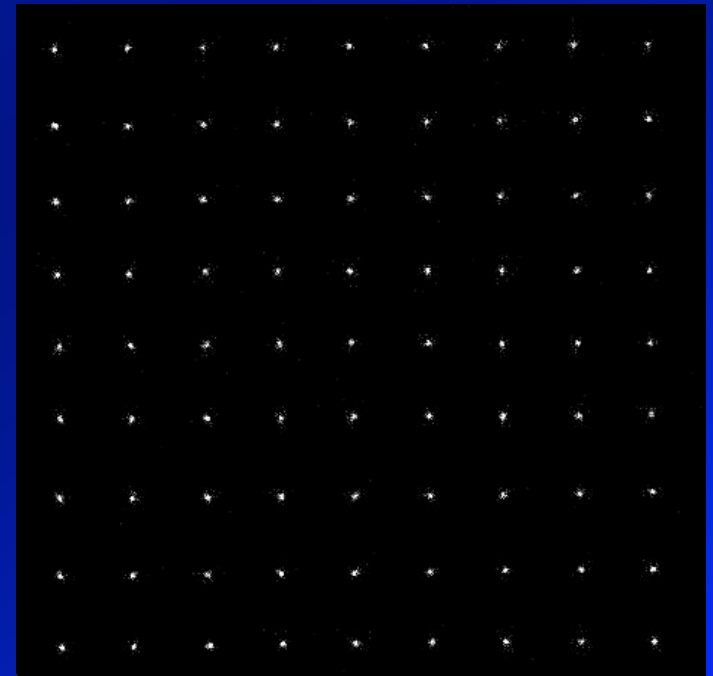




# *Imaging tests - resolution*



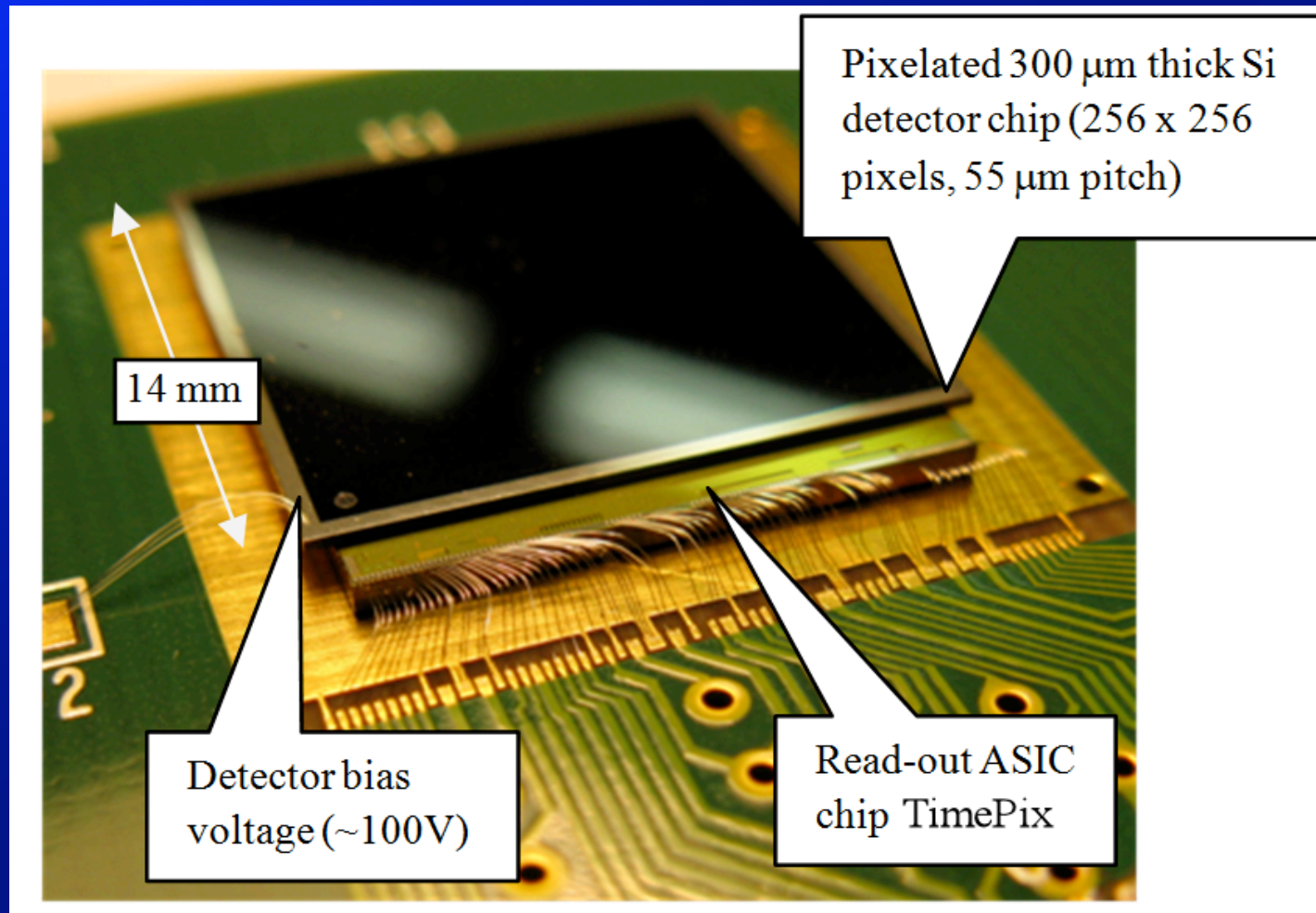
40 mm - 0.5 x 0.5 mm pinhole grid



Zoomed - 20  $\mu\text{m}$  FWHM avg.

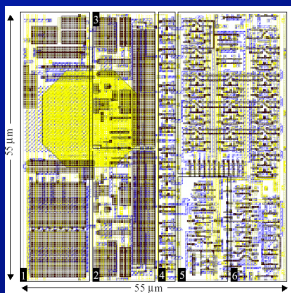
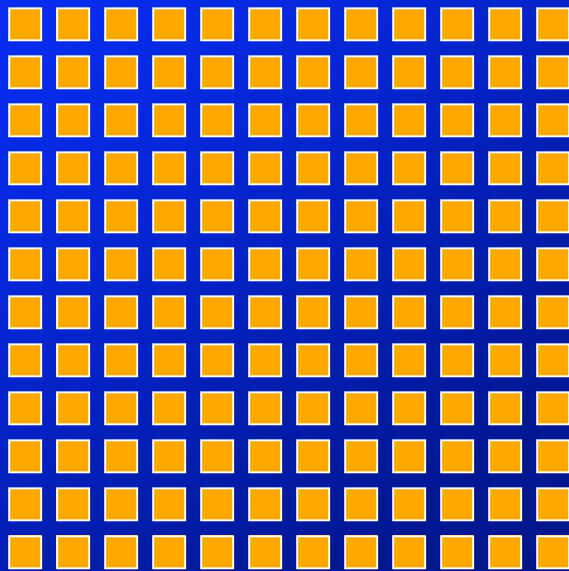


# Medipix/Timepix x-ray imager





# Medipix/Timepix ASIC readout



~ 500 transistors/pixel

- 256 x 256 array of 55  $\mu\text{m}$  pixels
- Integrates *counts*, not charge
- 100 kHz/pxl
- Frame rate: 1 kHz
- Low noise ( $100e^-$ ) = low gain operation ( $2\text{ ke}^-$ )
- GHz global count rate
- ~1 W watt/chip, abutable
- Developed at CERN



## Aside: Timepix readout of APD arrays?

Require gains of  $> 1000$  but  $< 200000$

Input pitch is  $55\mu\text{m}$

256 x 256 array

Can cool CMOS chip to 77K

(but 1W per chip is a high load)

Cannot bias pixels separately

Input takes holes or electrons,

(can sink up to  $10\text{nA/pxl}$ )





# Old WWII watch movie



66 MHz input rate



# Old WWII watch movie 2



*Bkgd*

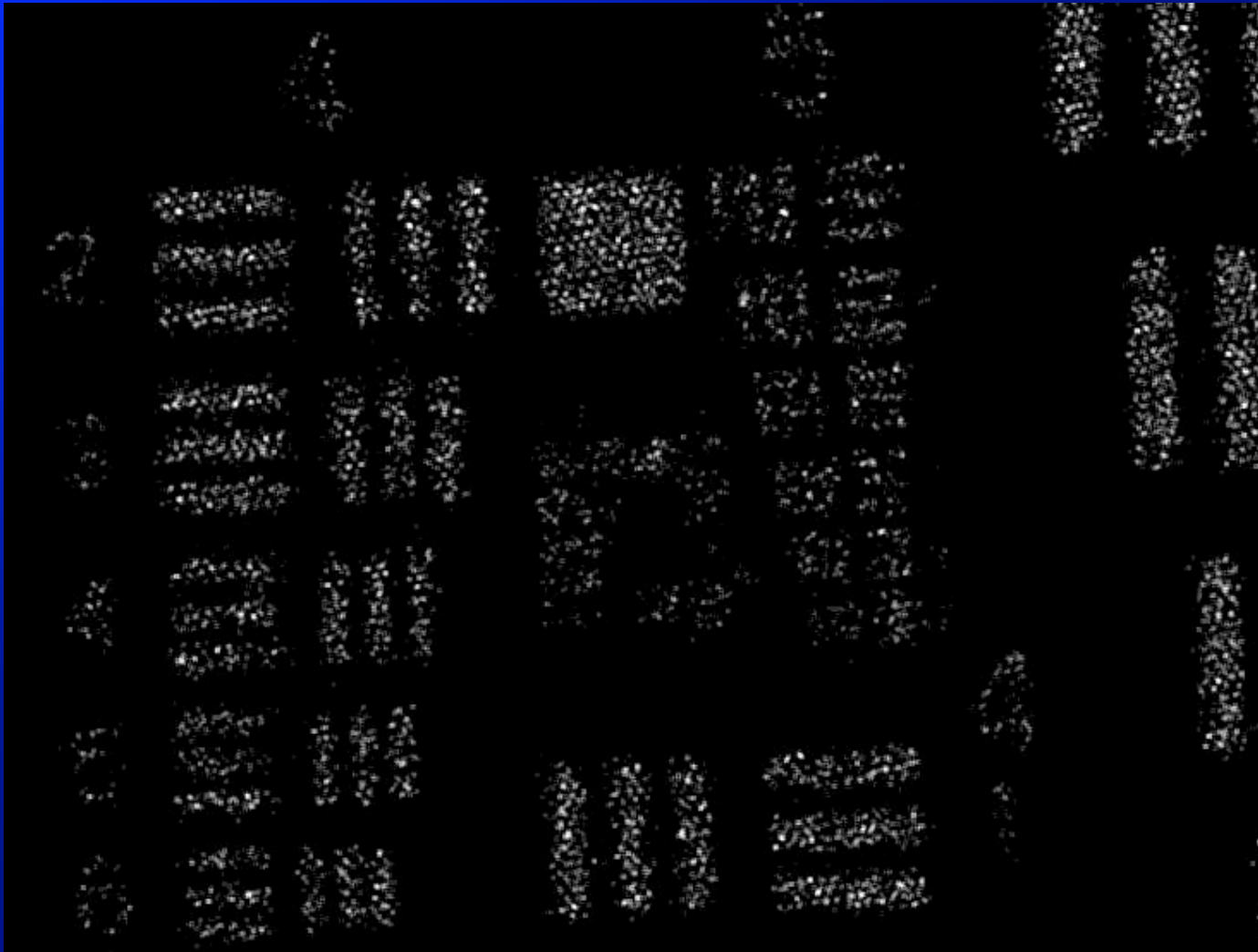
*.002 ct/pxl/s*

*Room Temp*

*(radium dial)*



## MCP pores (10 on 12 micron)





# Evolution of MCP detectors

	<u>1985</u>	<u>2010</u>	<u>Improvement</u>
Count rate (cps)	5000	$5 \times 10^6$	1000
Spatial Resolution ( $\mu\text{m}$ )	100	10	10
Lifetime (cts/mm)	$10^9$	$10^{12}$	1000
Optical QE (%)	~15	~30	2

*Limited by MCP technology!*

Size (mm)	100	100	1
Pore diameter ( $\mu\text{m}$ )	12	6	2
Lifetime (C/mm)	.01	.01	1





# Atomic layer deposition (ALD)

Allows engineering of surfaces after MCP pores are fabricated

- Optimize secondary electron coefficient for gain
- Optimize resistivity

Separates microchannel fabrication from surface preparation

- MCPs can be made from any glass, alumina, plastic
- Lithographic techniques for larger arrays of pores

Photocathodes deposited on MCP(?)

- Passivated ( $\text{Al}_2\text{O}_3$ ) surfaces
- Higher QE of opaque photocathodes



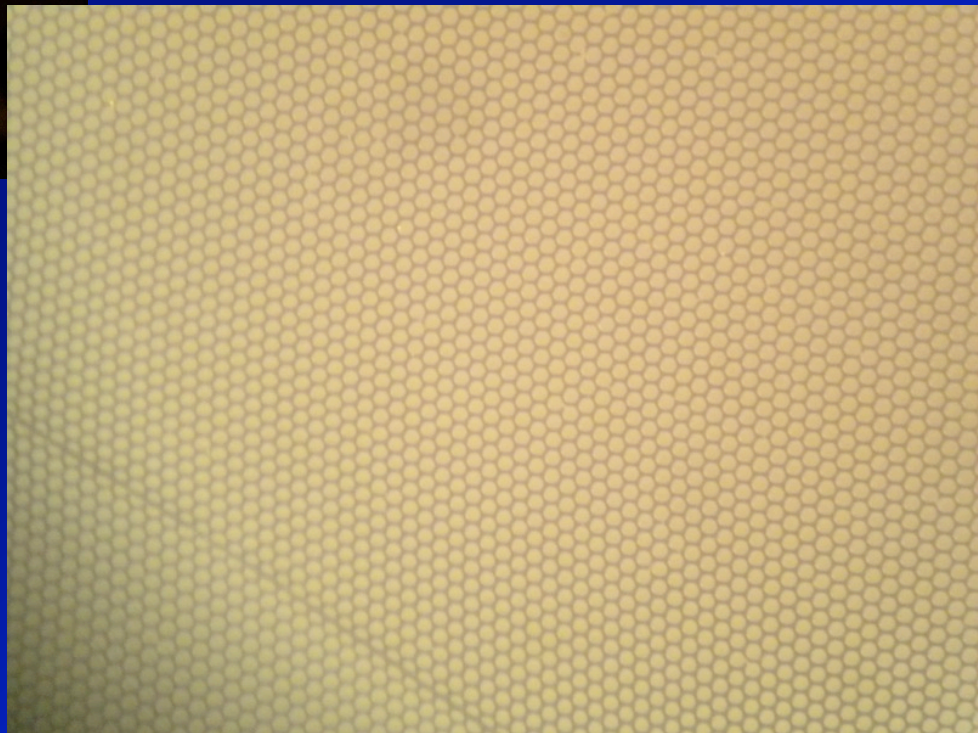
# Borosilicate glass MCP



33 mm substrate from INCOM

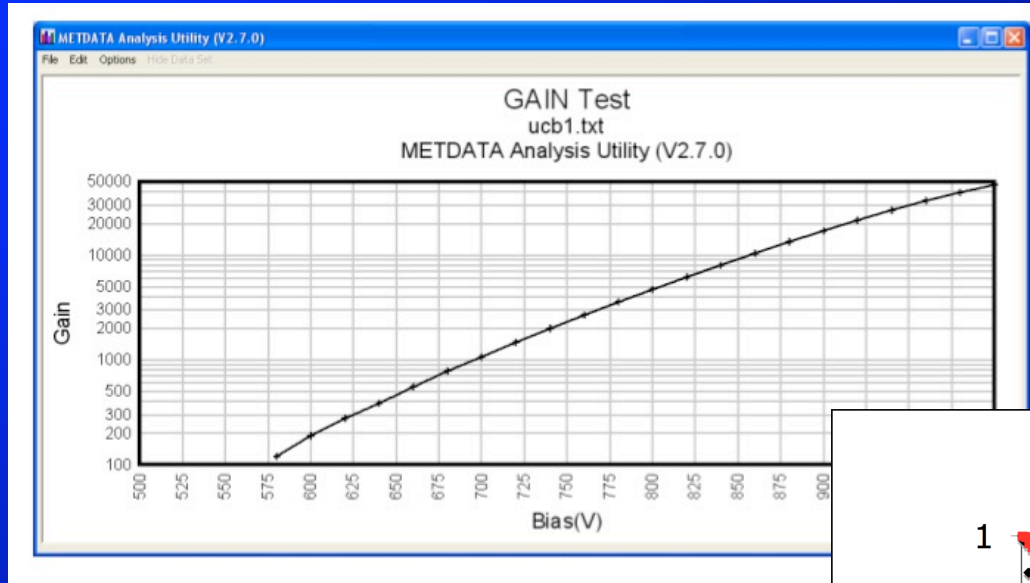
Resistive, SEC, and conductive coatings by Arradance

Microscope image showing  
40  $\mu\text{m}$  holes in hexagonal  
array with L/d ratio of 40

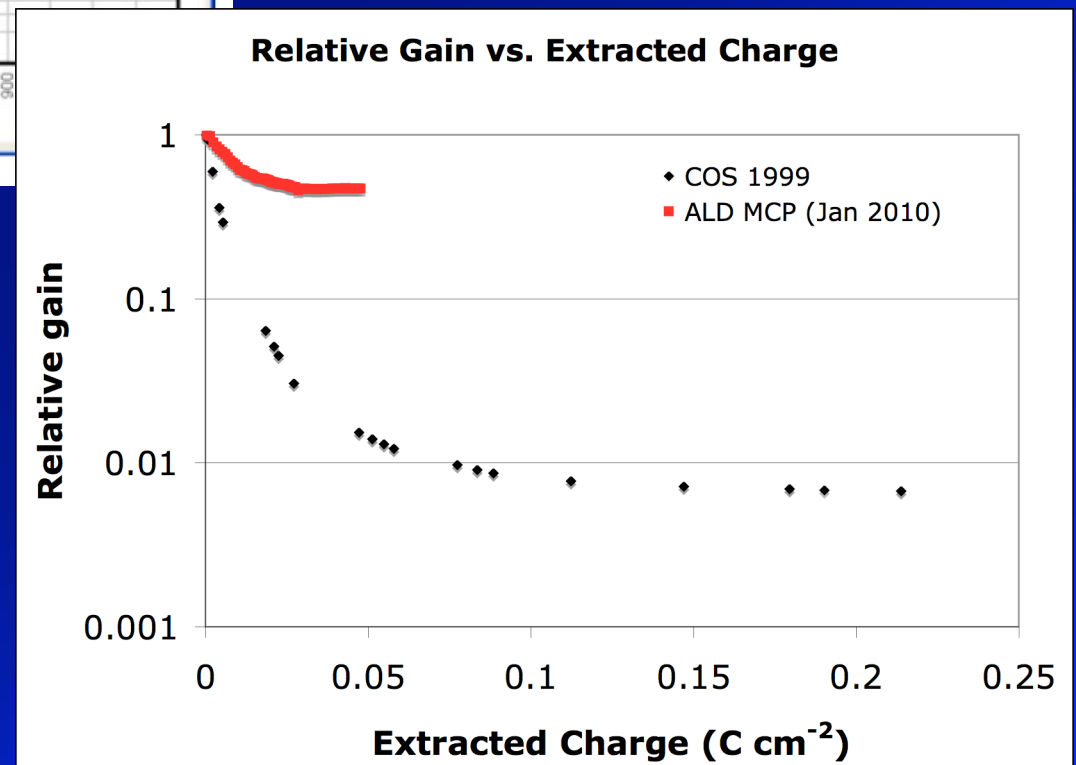




# Results from ALD coated Incom glass



ALD coating done by  
Arradiance Inc.  
([www.arradiance.com](http://www.arradiance.com))

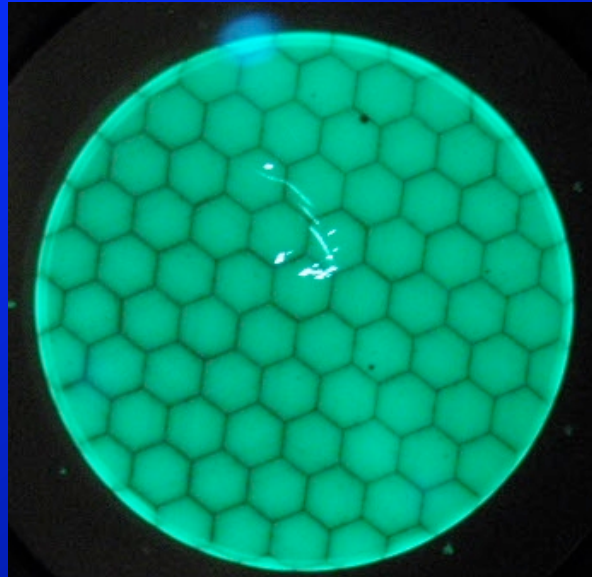






# SSL-UCB, ALD/Incom MCP Test

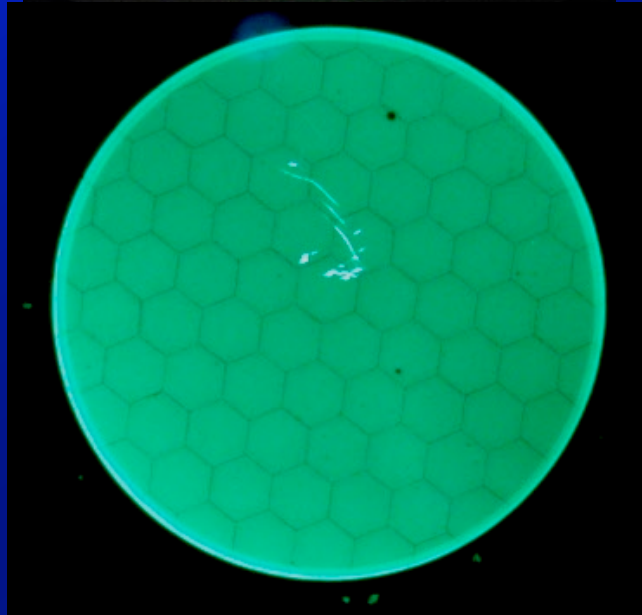
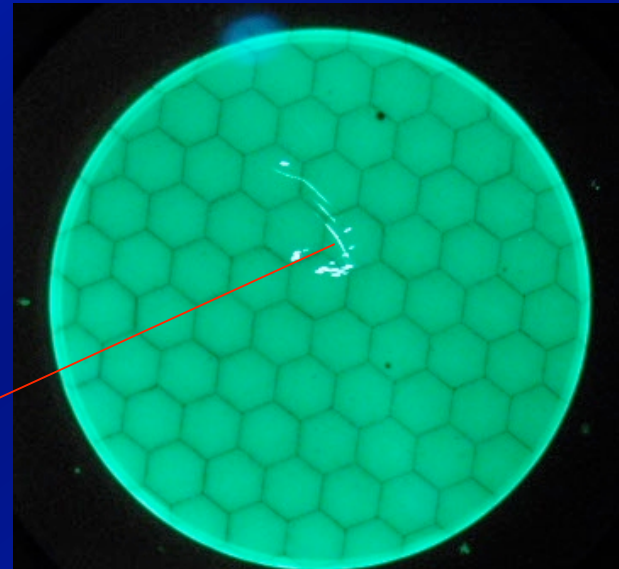
It works!



700v

800v

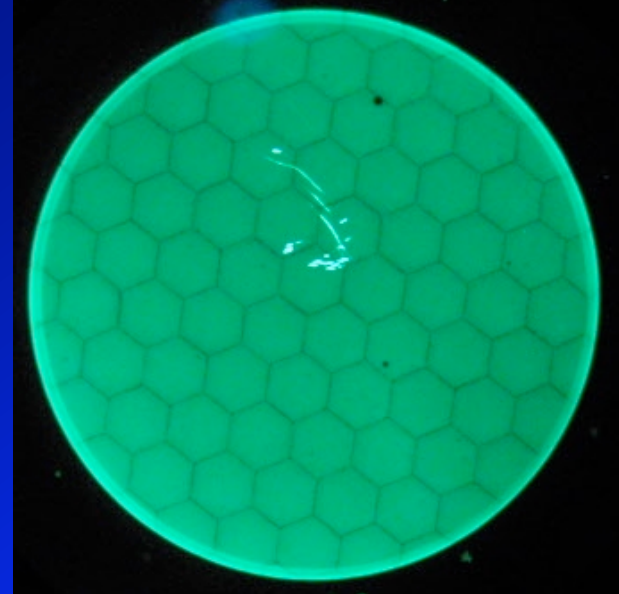
Bright scratch  
on phosphor



1000v

UV light

900v







# Psec Timing Project

(P.I Henry Frisch, U. Chicago, Argonne)

Large area (8"x8") MCP image tubes for  
Cherenkov arrays and PET detectors

Delayline readout, specialized ASICs

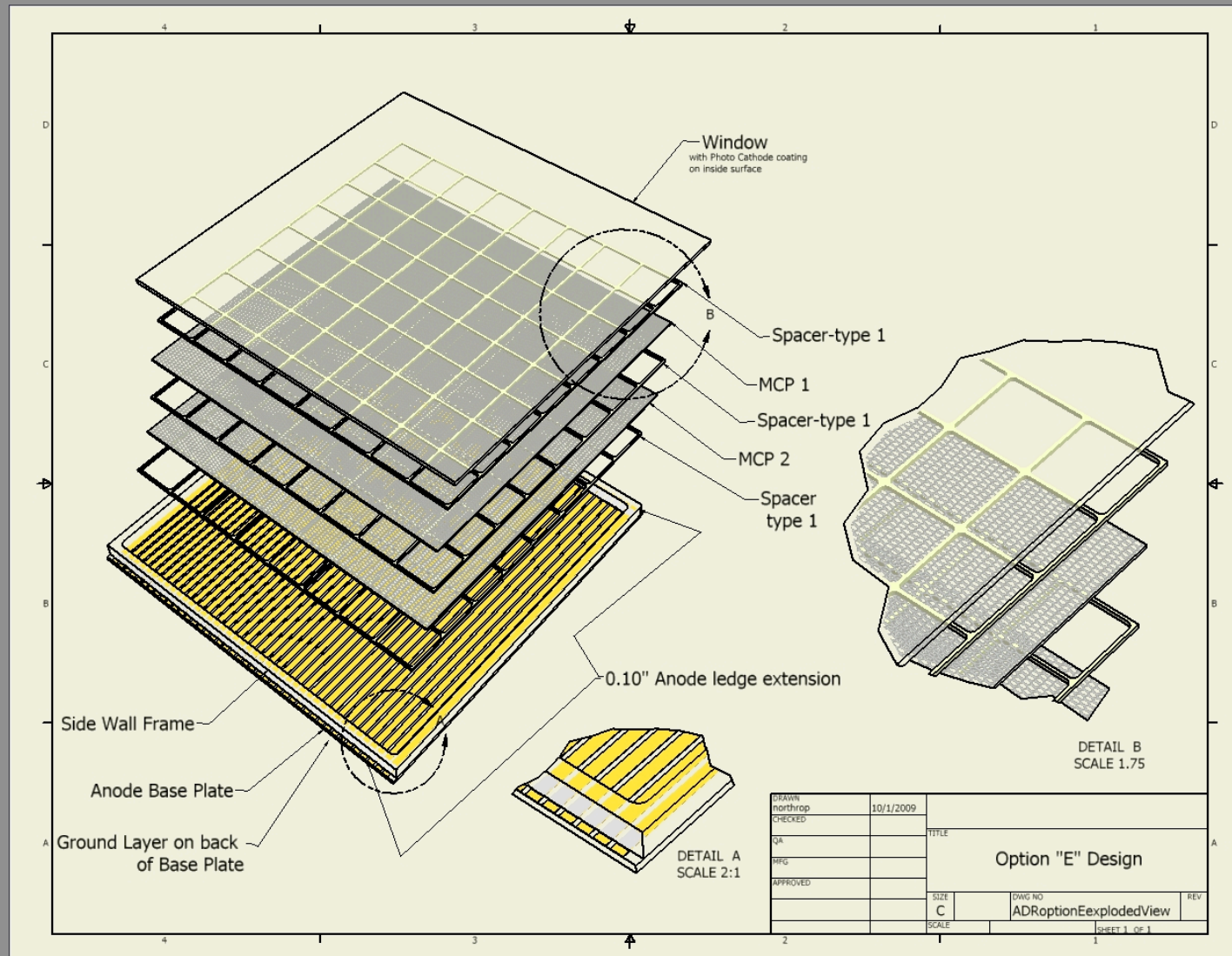
Pulse timing to 1 ps (multiple photon)

Thousands of tubes

Inexpensive design



# 8 inch image tube





# Concluding Remarks

*MCP detectors “in the field” demonstrate what can be done with an imaging photon counting detectors.*

*They are prevalent in the UV because of their historical QE advantage.*

*Modern electronic techniques are expanding their niche applications*

*New developments in optical/near IR photocathodes are pushing this technology to longer wavelengths.*

*Other known weaknesses (lifetime, cost, etc.) are being addressed with new materials and methods.*

*There is still life in this old war-horse!*





# Thank You

I would like to acknowledge the support of the  
*National Science Foundation*  
*National Institutes of Health*  
*National Institutes of Standards and Technology*  
*National Aeronautics and Space Administration*  
under various contract and grant numbers.